



Observation of the Bottomonium Ground State in the Decay $Y(3S) \rightarrow \gamma\eta_b$

B. Aubert,¹ M. Bona,¹ Y. Karyotakis,¹ J. P. Lees,¹ V. Poireau,¹ E. Prencipe,¹ X. Prudent,¹ V. Tisserand,¹ J. Garra Tico,² E. Grauges,² L. Lopez,³ A. Palano,³ M. Pappagallo,³ G. Eigen,⁴ B. Stugu,⁴ L. Sun,⁴ G. S. Abrams,⁵ M. Battaglia,⁵ D. N. Brown,⁵ R. N. Cahn,⁵ R. G. Jacobsen,⁵ L. T. Kerth,⁵ Yu. G. Kolomensky,⁵ G. Lynch,⁵ I. L. Osipenko,⁵ M. T. Ronan,^{5,*} K. Tackmann,⁵ T. Tanabe,⁵ C. M. Hawkes,⁶ N. Soni,⁶ A. T. Watson,⁶ H. Koch,⁷ T. Schroeder,⁷ D. Walker,⁸ D. J. Asgeirsson,⁹ B. G. Fulsom,⁹ C. Hearty,⁹ T. S. Mattison,⁹ J. A. McKenna,⁹ M. Barrett,¹⁰ A. Khan,¹⁰ V. E. Blinov,¹¹ A. D. Bukin,¹¹ A. R. Buzykaev,¹¹ V. P. Druzhinin,¹¹ V. B. Golubev,¹¹ A. P. Onuchin,¹¹ S. I. Serednyakov,¹¹ Yu. I. Skovpen,¹¹ E. P. Solodov,¹¹ K. Yu. Todyshev,¹¹ M. Bondioli,¹² S. Curry,¹² I. Eschrich,¹² D. Kirkby,¹² A. J. Lankford,¹² P. Lund,¹² M. Mandelkern,¹² E. C. Martin,¹² D. P. Stoker,¹² S. Abachi,¹³ C. Buchanan,¹³ J. W. Gary,¹⁴ F. Liu,¹⁴ O. Long,¹⁴ B. C. Shen,^{14,*} G. M. Vitug,¹⁴ Z. Yasin,¹⁴ L. Zhang,¹⁴ V. Sharma,¹⁵ C. Campagnari,¹⁶ T. M. Hong,¹⁶ D. Kovalskyi,¹⁶ M. A. Mazur,¹⁶ J. D. Richman,¹⁶ T. W. Beck,¹⁷ A. M. Eisner,¹⁷ C. J. Flacco,¹⁷ C. A. Heusch,¹⁷ J. Kroseberg,¹⁷ W. S. Lockman,¹⁷ A. J. Martinez,¹⁷ T. Schalk,¹⁷ B. A. Schumm,¹⁷ A. Seiden,¹⁷ M. G. Wilson,¹⁷ L. O. Winstrom,¹⁷ C. H. Cheng,¹⁸ D. A. Doll,¹⁸ B. Echenard,¹⁸ F. Fang,¹⁸ D. G. Hitlin,¹⁸ I. Narsky,¹⁸ T. Piatenko,¹⁸ F. C. Porter,¹⁸ R. Andreassen,¹⁹ G. Mancinelli,¹⁹ B. T. Meadows,¹⁹ K. Mishra,¹⁹ M. D. Sokoloff,¹⁹ P. C. Bloom,²⁰ W. T. Ford,²⁰ A. Gaz,²⁰ J. F. Hirschauer,²⁰ M. Nagel,²⁰ U. Nauenberg,²⁰ J. G. Smith,²⁰ K. A. Ulmer,²⁰ S. R. Wagner,²⁰ R. Ayad,^{21,†} A. Soffer,^{21,‡} W. H. Toki,²¹ R. J. Wilson,²¹ D. D. Altenburg,²² E. Feltres,²² A. Hauke,²² H. Jasper,²² M. Karbach,²² J. Merkel,²² A. Petzold,²² B. Spaan,²² K. Wacker,²² M. J. Kobel,²³ W. F. Mader,²³ R. Nogowski,²³ K. R. Schubert,²³ R. Schwier,²³ A. Volk,²³ D. Bernard,²⁴ G. R. Bonneaud,²⁴ E. Latour,²⁴ M. Verderi,²⁴ P. J. Clark,²⁵ S. Playfer,²⁵ J. E. Watson,²⁵ M. Andreotti,²⁶ D. Bettoni,²⁶ C. Bozzi,²⁶ R. Calabrese,²⁶ A. Cecchi,²⁶ G. Cibinetto,²⁶ P. Franchini,²⁶ E. Luppi,²⁶ M. Negrini,²⁶ A. Petrella,²⁶ L. Piemontese,²⁶ V. Santoro,²⁶ R. Baldini-Ferroli,²⁷ A. Calcaterra,²⁷ R. de Sangro,²⁷ G. Finocchiaro,²⁷ S. Pacetti,²⁷ P. Patteri,²⁷ I. M. Peruzzi,^{27,§} M. Piccolo,²⁷ M. Rama,²⁷ A. Zallo,²⁷ A. Buzzo,²⁸ R. Contri,²⁸ M. Lo Vetere,²⁸ M. M. Macri,²⁸ M. R. Monge,²⁸ S. Passaggio,²⁸ C. Patrignani,²⁸ E. Robutti,²⁸ A. Santroni,²⁸ S. Tosi,²⁸ K. S. Chaisanguanthum,²⁹ M. Morii,²⁹ A. Adametz,³⁰ J. Marks,³⁰ S. Schenk,³⁰ U. Uwer,³⁰ V. Klose,³¹ H. M. Lacker,³¹ D. J. Bard,³² P. D. Dauncey,³² J. A. Nash,³² M. Tibbetts,³² P. K. Behera,³³ X. Chai,³³ M. J. Charles,³³ U. Mallik,³³ J. Cochran,³⁴ H. B. Crawley,³⁴ L. Dong,³⁴ W. T. Meyer,³⁴ S. Prell,³⁴ E. I. Rosenberg,³⁴ A. E. Rubin,³⁴ Y. Y. Gao,³⁵ A. V. Gritsan,³⁵ Z. J. Guo,³⁵ C. K. Lae,³⁵ N. Arnaud,³⁶ J. Béquilleux,³⁶ A. D'Orazio,³⁶ M. Davier,³⁶ J. Firmino da Costa,³⁶ G. Grosdidier,³⁶ A. Höcker,³⁶ V. Lepeltier,³⁶ F. Le Diberder,³⁶ A. M. Lutz,³⁶ S. Pruvot,³⁶ P. Roudeau,³⁶ M. H. Schune,³⁶ J. Serrano,³⁶ V. Sordini,^{36,||} A. Stocchi,³⁶ G. Wormser,³⁶ D. J. Lange,³⁷ D. M. Wright,³⁷ I. Bingham,³⁸ J. P. Burke,³⁸ C. A. Chavez,³⁸ J. R. Fry,³⁸ E. Gabathuler,³⁸ R. Gamet,³⁸ D. E. Hutchcroft,³⁸ D. J. Payne,³⁸ C. Touramanis,³⁸ A. J. Bevan,³⁹ C. K. Clarke,³⁹ K. A. George,³⁹ F. Di Lodovico,³⁹ R. Sacco,³⁹ M. Sigamani,³⁹ G. Cowan,⁴⁰ H. U. Flaecher,⁴⁰ D. A. Hopkins,⁴⁰ S. Paramesvaran,⁴⁰ F. Salvatore,⁴⁰ A. C. Wren,⁴⁰ D. N. Brown,⁴¹ C. L. Davis,⁴¹ A. G. Denig,⁴² M. Fritsch,⁴² W. Gradl,⁴² G. Schott,⁴² K. E. Alwyn,⁴³ D. Bailey,⁴³ R. J. Barlow,⁴³ Y. M. Chia,⁴³ C. L. Edgar,⁴³ G. Jackson,⁴³ G. D. Lafferty,⁴³ T. J. West,⁴³ J. I. Yi,⁴³ J. Anderson,⁴⁴ C. Chen,⁴⁴ A. Jawahery,⁴⁴ D. A. Roberts,⁴⁴ G. Simi,⁴⁴ J. M. Tuggle,⁴⁴ C. Dallapiccola,⁴⁵ X. Li,⁴⁵ E. Salvati,⁴⁵ S. Saremi,⁴⁵ R. Cowan,⁴⁶ D. Dujmic,⁴⁶ P. H. Fisher,⁴⁶ G. Sciolla,⁴⁶ M. Spitznagel,⁴⁶ F. Taylor,⁴⁶ R. K. Yamamoto,⁴⁶ M. Zhao,⁴⁶ P. M. Patel,⁴⁷ S. H. Robertson,⁴⁷ A. Lazzaro,⁴⁸ V. Lombardo,⁴⁸ F. Palombo,⁴⁸ J. M. Bauer,⁴⁹ L. Cremaldi,⁴⁹ R. Godang,^{49,¶} R. Kroeger,⁴⁹ D. A. Sanders,⁴⁹ D. J. Summers,⁴⁹ H. W. Zhao,⁴⁹ M. Simard,⁵⁰ P. Taras,⁵⁰ F. B. Viaud,⁵⁰ H. Nicholson,⁵¹ G. De Nardo,⁵² L. Lista,⁵² D. Monorchio,⁵² G. Onorato,⁵² C. Sciacca,⁵² G. Raven,⁵³ H. L. Snoek,⁵³ C. P. Jessop,⁵⁴ K. J. Knoepfel,⁵⁴ J. M. LoSecco,⁵⁴ W. F. Wang,⁵⁴ G. Benelli,⁵⁵ L. A. Corwin,⁵⁵ K. Honscheid,⁵⁵ H. Kagan,⁵⁵ R. Kass,⁵⁵ J. P. Morris,⁵⁵ A. M. Rahimi,⁵⁵ J. J. Regensburger,⁵⁵ S. J. Sekula,⁵⁵ Q. K. Wong,⁵⁵ N. L. Blount,⁵⁶ J. Brau,⁵⁶ R. Frey,⁵⁶ O. Igonkina,⁵⁶ J. A. Kolb,⁵⁶ M. Lu,⁵⁶ R. Rahmat,⁵⁶ N. B. Sinev,⁵⁶ D. Strom,⁵⁶ J. Strube,⁵⁶ E. Torrence,⁵⁶ G. Castelli,⁵⁷ N. Gagliardi,⁵⁷ M. Margoni,⁵⁷ M. Morandin,⁵⁷ M. Posocco,⁵⁷ M. Rotondo,⁵⁷ F. Simonetto,⁵⁷ R. Stroili,⁵⁷ C. Voci,⁵⁷ P. del Amo Sanchez,⁵⁸ E. Ben-Haim,⁵⁸ H. Briand,⁵⁸ G. Calderini,⁵⁸ J. Chauveau,⁵⁸ P. David,⁵⁸ L. Del Buono,⁵⁸ O. Hamon,⁵⁸ Ph. Leruste,⁵⁸ J. Ocariz,⁵⁸ A. Perez,⁵⁸ J. Prendki,⁵⁸ S. Sitt,⁵⁸ L. Gladney,⁵⁹ M. Biasini,⁶⁰ R. Covarelli,⁶⁰ E. Manoni,⁶⁰ C. Angelini,⁶¹ G. Batignani,⁶¹ S. Bettarini,⁶¹ M. Carpinelli,^{61,*} A. Cervelli,⁶¹ F. Forti,⁶¹ M. A. Giorgi,⁶¹ A. Lusiani,⁶¹ G. Marchiori,⁶¹ M. Morganti,⁶¹ N. Neri,⁶¹ E. Paoloni,⁶¹ G. Rizzo,⁶¹ J. J. Walsh,⁶¹ D. Lopes Pegna,⁶² C. Lu,⁶² J. Olsen,⁶² A. J. S. Smith,⁶² A. V. Telnov,⁶² F. Anulli,⁶³ E. Baracchini,⁶³ G. Cavoto,⁶³ D. del Re,⁶³ E. Di Marco,⁶³ R. Faccini,⁶³ F. Ferrarotto,⁶³ F. Ferroni,⁶³ M. Gaspero,⁶³ P. D. Jackson,⁶³ L. Li Gioi,⁶³ M. A. Mazzoni,⁶³ S. Morganti,⁶³ G. Piredda,⁶³ F. Polci,⁶³ F. Renga,⁶³ C. Voena,⁶³ M. Ebert,⁶⁴ T. Hartmann,⁶⁴ H. Schröder,⁶⁴ R. Waldi,⁶⁴ T. Adye,⁶⁵

B. Franek,⁶⁵ E. O. Olaiya,⁶⁵ F. F. Wilson,⁶⁵ S. Emery,⁶⁶ M. Escalier,⁶⁶ L. Esteve,⁶⁶ S. F. Ganzhur,⁶⁶ G. Hamel de Monchenault,⁶⁶ W. Kozanecki,⁶⁶ G. Vasseur,⁶⁶ Ch. Yèche,⁶⁶ M. Zito,⁶⁶ X. R. Chen,⁶⁷ H. Liu,⁶⁷ W. Park,⁶⁷ M. V. Purohit,⁶⁷ R. M. White,⁶⁷ J. R. Wilson,⁶⁷ M. T. Allen,⁶⁸ D. Aston,⁶⁸ R. Bartoldus,⁶⁸ P. Bechtle,⁶⁸ J. F. Benitez,⁶⁸ K. Bertsche,⁶⁸ Y. Cai,⁶⁸ R. Cenci,⁶⁸ J. P. Coleman,⁶⁸ M. R. Convery,⁶⁸ F. J. Decker,⁶⁸ J. C. Dingfelder,⁶⁸ J. Dorfan,⁶⁸ G. P. Dubois-Felsmann,⁶⁸ W. Dunwoodie,⁶⁸ S. Ecklund,⁶⁸ R. Erickson,⁶⁸ R. C. Field,⁶⁸ A. Fisher,⁶⁸ J. Fox,⁶⁸ A. M. Gabareen,⁶⁸ S. J. Gowdy,⁶⁸ M. T. Graham,⁶⁸ P. Grenier,⁶⁸ C. Hast,⁶⁸ W. R. Innes,⁶⁸ R. Iverson,⁶⁸ J. Kaminski,⁶⁸ M. H. Kelsey,⁶⁸ H. Kim,⁶⁸ P. Kim,⁶⁸ M. L. Kocian,⁶⁸ A. Kulikov,⁶⁸ D. W. G. S. Leith,⁶⁸ S. Li,⁶⁸ B. Lindquist,⁶⁸ S. Luitz,⁶⁸ V. Luth,⁶⁸ H. L. Lynch,⁶⁸ D. B. MacFarlane,⁶⁸ H. Marsiske,⁶⁸ R. Messner,⁶⁸ D. R. Muller,⁶⁸ H. Neal,⁶⁸ S. Nelson,⁶⁸ A. Novokhatski,⁶⁸ C. P. O'Grady,⁶⁸ I. Ofte,⁶⁸ A. Perazzo,⁶⁸ M. Perl,⁶⁸ B. N. Ratcliff,⁶⁸ C. Rivetta,⁶⁸ A. Roodman,⁶⁸ A. A. Salnikov,⁶⁸ R. H. Schindler,⁶⁸ J. Schwiening,⁶⁸ J. Seeman,⁶⁸ A. Snyder,⁶⁸ D. Su,⁶⁸ M. K. Sullivan,⁶⁸ K. Suzuki,⁶⁸ S. K. Swain,⁶⁸ J. M. Thompson,⁶⁸ J. Va'vra,⁶⁸ D. Van Winkle,⁶⁸ A. P. Wagner,⁶⁸ M. Weaver,⁶⁸ C. A. West,⁶⁸ U. Wienands,⁶⁸ W. J. Wisniewski,⁶⁸ M. Wittgen,⁶⁸ W. Wittmer,⁶⁸ D. H. Wright,⁶⁸ H. W. Wulsin,⁶⁸ Y. Yan,⁶⁸ A. K. Yarritu,⁶⁸ K. Yi,⁶⁸ G. Yocky,⁶⁸ C. C. Young,⁶⁸ V. Ziegler,⁶⁸ P. R. Burchat,⁶⁹ A. J. Edwards,⁶⁹ S. A. Majewski,⁶⁹ T. S. Miyashita,⁶⁹ B. A. Petersen,⁶⁹ L. Wilden,⁶⁹ S. Ahmed,⁷⁰ M. S. Alam,⁷⁰ J. A. Ernst,⁷⁰ B. Pan,⁷⁰ M. A. Saeed,⁷⁰ S. B. Zain,⁷⁰ S. M. Spanier,⁷¹ B. J. Wogslund,⁷¹ R. Eckmann,⁷² J. L. Ritchie,⁷² A. M. Ruland,⁷² C. J. Schilling,⁷² R. F. Schwitters,⁷² B. W. Drummond,⁷³ J. M. Izen,⁷³ X. C. Lou,⁷³ F. Bianchi,⁷⁴ D. Gamba,⁷⁴ M. Pelliccioni,⁷⁴ M. Bomben,⁷⁵ L. Bosisio,⁷⁵ C. Cartaro,⁷⁵ G. Della Ricca,⁷⁵ L. Lanceri,⁷⁵ L. Vitale,⁷⁵ V. Azzolini,⁷⁶ N. Lopez-March,⁷⁶ F. Martinez-Vidal,⁷⁶ D. A. Milanes,⁷⁶ A. Oyanguen,⁷⁶ J. Albert,⁷⁷ Sw. Banerjee,⁷⁷ B. Bhuyan,⁷⁷ H. H. F. Choi,⁷⁷ K. Hamano,⁷⁷ R. Kowalewski,⁷⁷ M. J. Lewczuk,⁷⁷ I. M. Nugent,⁷⁷ J. M. Roney,⁷⁷ R. J. Sobie,⁷⁷ T. J. Gershon,⁷⁸ P. F. Harrison,⁷⁸ J. Ilic,⁷⁸ T. E. Latham,⁷⁸ G. B. Mohanty,⁷⁸ H. R. Band,⁷⁹ X. Chen,⁷⁹ S. Dasu,⁷⁹ K. T. Flood,⁷⁹ Y. Pan,⁷⁹ M. Pierini,⁷⁹ R. Prepost,⁷⁹ C. O. Vuosalo,⁷⁹ and S. L. Wu⁷⁹

(BABAR Collaboration)

¹Laboratoire de Physique des Particules, IN2P3/CNRS et Université de Savoie, F-74941 Annecy-Le-Vieux, France

²Universitat de Barcelona, Facultat de Física, Departament ECM, E-08028 Barcelona, Spain

³INFN Sezione di Bari; Dipartimento di Fisica, Università di Bari, I-70126 Bari, Italy

⁴University of Bergen, Institute of Physics, N-5007 Bergen, Norway

⁵Lawrence Berkeley National Laboratory and University of California, Berkeley, California 94720, USA

⁶University of Birmingham, Birmingham, B15 2TT, United Kingdom

⁷Ruhr Universität Bochum, Institut für Experimentalphysik 1, D-44780 Bochum, Germany

⁸University of Bristol, Bristol BS8 1TL, United Kingdom

⁹University of British Columbia, Vancouver, British Columbia, Canada V6T 1Z1

¹⁰Brunel University, Uxbridge, Middlesex UB8 3PH, United Kingdom

¹¹Budker Institute of Nuclear Physics, Novosibirsk 630090, Russia

¹²University of California at Irvine, Irvine, California 92697, USA

¹³University of California at Los Angeles, Los Angeles, California 90024, USA

¹⁴University of California at Riverside, Riverside, California 92521, USA

¹⁵University of California at San Diego, La Jolla, California 92093, USA

¹⁶University of California at Santa Barbara, Santa Barbara, California 93106, USA

¹⁷University of California at Santa Cruz, Institute for Particle Physics, Santa Cruz, California 95064, USA

¹⁸California Institute of Technology, Pasadena, California 91125, USA

¹⁹University of Cincinnati, Cincinnati, Ohio 45221, USA

²⁰University of Colorado, Boulder, Colorado 80309, USA

²¹Colorado State University, Fort Collins, Colorado 80523, USA

²²Technische Universität Dortmund, Fakultät Physik, D-44221 Dortmund, Germany

²³Technische Universität Dresden, Institut für Kern- und Teilchenphysik, D-01062 Dresden, Germany

²⁴Laboratoire Leprince-Ringuet, CNRS/IN2P3, Ecole Polytechnique, F-91128 Palaiseau, France

²⁵University of Edinburgh, Edinburgh EH9 3JZ, United Kingdom

²⁶INFN Sezione di Ferrara; Dipartimento di Fisica, Università di Ferrara, I-44100 Ferrara, Italy

²⁷INFN Laboratori Nazionali di Frascati, I-00044 Frascati, Italy

²⁸INFN Sezione di Genova; Dipartimento di Fisica, Università di Genova, I-16146 Genova, Italy

²⁹Harvard University, Cambridge, Massachusetts 02138, USA

³⁰Universität Heidelberg, Physikalisches Institut, Philosophenweg 12, D-69120 Heidelberg, Germany

³¹Humboldt-Universität zu Berlin, Institut für Physik, Newtonstr. 15, D-12489 Berlin, Germany

³²Imperial College London, London, SW7 2AZ, United Kingdom

- ³³University of Iowa, Iowa City, Iowa 52242, USA
³⁴Iowa State University, Ames, Iowa 50011-3160, USA
³⁵Johns Hopkins University, Baltimore, Maryland 21218, USA
³⁶Laboratoire de l'Accélérateur Linéaire, IN2P3/CNRS et Université Paris-Sud 11, Centre Scientifique d'Orsay, B. P. 34, F-91898 Orsay Cedex, France
³⁷Lawrence Livermore National Laboratory, Livermore, California 94550, USA
³⁸University of Liverpool, Liverpool L69 7ZE, United Kingdom
³⁹Queen Mary, University of London, London, E1 4NS, United Kingdom
⁴⁰University of London, Royal Holloway and Bedford New College, Egham, Surrey TW20 0EX, United Kingdom
⁴¹University of Louisville, Louisville, Kentucky 40292, USA
⁴²Johannes Gutenberg-Universität Mainz, Institut für Kernphysik, D-55099 Mainz, Germany
⁴³University of Manchester, Manchester M13 9PL, United Kingdom
⁴⁴University of Maryland, College Park, Maryland 20742, USA
⁴⁵University of Massachusetts, Amherst, Massachusetts 01003, USA
⁴⁶Massachusetts Institute of Technology, Laboratory for Nuclear Science, Cambridge, Massachusetts 02139, USA
⁴⁷McGill University, Montréal, Québec, Canada H3A 2T8
⁴⁸INFN Sezione di Milano; Dipartimento di Fisica, Università di Milano, I-20133 Milano, Italy
⁴⁹University of Mississippi, University, Mississippi 38677, USA
⁵⁰Université de Montréal, Physique des Particules, Montréal, Québec, Canada H3C 3J7
⁵¹Mount Holyoke College, South Hadley, Massachusetts 01075, USA
⁵²INFN Sezione di Napoli; Dipartimento di Scienze Fisiche, Università di Napoli Federico II, I-80126 Napoli, Italy
⁵³NIKHEF, National Institute for Nuclear Physics and High Energy Physics, NL-1009 DB Amsterdam, The Netherlands
⁵⁴University of Notre Dame, Notre Dame, Indiana 46556, USA
⁵⁵Ohio State University, Columbus, Ohio 43210, USA
⁵⁶University of Oregon, Eugene, Oregon 97403, USA
⁵⁷INFN Sezione di Padova; Dipartimento di Fisica, Università di Padova, I-35131 Padova, Italy
⁵⁸Laboratoire de Physique Nucléaire et de Hautes Energies, IN2P3/CNRS, Université Pierre et Marie Curie-Paris6, Université Denis Diderot-Paris7, F-75252 Paris, France
⁵⁹University of Pennsylvania, Philadelphia, Pennsylvania 19104, USA
⁶⁰INFN Sezione di Perugia; Dipartimento di Fisica, Università di Perugia, I-06100 Perugia, Italy
⁶¹INFN Sezione di Pisa; Dipartimento di Fisica, Università di Pisa; Scuola Normale Superiore di Pisa, I-56127 Pisa, Italy
⁶²Princeton University, Princeton, New Jersey 08544, USA
⁶³INFN Sezione di Roma; Dipartimento di Fisica, Università di Roma La Sapienza, I-00185 Roma, Italy
⁶⁴Universität Rostock, D-18051 Rostock, Germany
⁶⁵Rutherford Appleton Laboratory, Chilton, Didcot, Oxon, OX11 0QX, United Kingdom
⁶⁶CEA, Irfu, SPP, Centre de Saclay, F-91191 Gif-sur-Yvette, France
⁶⁷University of South Carolina, Columbia, South Carolina 29208, USA
⁶⁸Stanford Linear Accelerator Center, Stanford, California 94309, USA
⁶⁹Stanford University, Stanford, California 94305-4060, USA
⁷⁰State University of New York, Albany, New York 12222, USA
⁷¹University of Tennessee, Knoxville, Tennessee 37996, USA
⁷²University of Texas at Austin, Austin, Texas 78712, USA
⁷³University of Texas at Dallas, Richardson, Texas 75083, USA
⁷⁴INFN Sezione di Torino; Dipartimento di Fisica Sperimentale, Università di Torino, I-10125 Torino, Italy
⁷⁵INFN Sezione di Trieste; Dipartimento di Fisica, Università di Trieste, I-34127 Trieste, Italy
⁷⁶IFIC, Universitat de Valencia-CSIC, E-46071 Valencia, Spain
⁷⁷University of Victoria, Victoria, British Columbia, Canada V8W 3P6
⁷⁸Department of Physics, University of Warwick, Coventry CV4 7AL, United Kingdom
⁷⁹University of Wisconsin, Madison, Wisconsin 53706, USA

(Received 7 July 2008; published 11 August 2008)

We report the results of a search for the bottomonium ground state $\eta_b(1S)$ in the photon energy spectrum with a sample of (109 ± 1) million of $Y(3S)$ recorded at the $Y(3S)$ energy with the BABAR detector at the PEP-II B factory at SLAC. We observe a peak in the photon energy spectrum at $E_\gamma = 921.2^{+2.1}_{-2.8}(\text{stat}) \pm 2.4(\text{syst})$ MeV with a significance of 10 standard deviations. We interpret the observed peak as being due to monochromatic photons from the radiative transition $Y(3S) \rightarrow \gamma\eta_b(1S)$. This photon energy corresponds to an $\eta_b(1S)$ mass of $9388.9^{+3.1}_{-2.3}(\text{stat}) \pm 2.7(\text{syst})$ MeV/ c^2 . The hyperfine $Y(1S)$ - $\eta_b(1S)$ mass splitting is $71.4^{+2.3}_{-3.1}(\text{stat}) \pm 2.7(\text{syst})$ MeV/ c^2 . The branching fraction for this radiative $Y(3S)$ decay is estimated to be $[4.8 \pm 0.5(\text{stat}) \pm 1.2(\text{syst})] \times 10^{-4}$.

Thirty years after the discovery of the narrow $Y(nS)$ resonances [1], no evidence has been reported for the spin-singlet pseudoscalar partners $\eta_b(nS)$ of these states. Measurement of the hyperfine mass splittings between the triplet and singlet states in quarkonium systems is of key importance in understanding the role of spin-spin interactions in quarkonium models and in testing QCD calculations [2]. Theoretical estimates of the mass splitting between the $1S$ singlet and triplet states vary from 36 to 100 MeV/ c^2 [3].

In this Letter, we report the observation of the radiative transition $Y(3S) \rightarrow \gamma \eta_b(1S)$, where the $\eta_b(1S)$, hereafter referred to as the η_b , is the pseudoscalar partner of the triplet state $Y(1S)$, and corresponds to the ground state of the bottomonium system. Theoretical predictions of the decay branching fraction range from 1 to 20×10^{-4} [3], where the unknown η_b mass is a major source of the uncertainties. The current limit from the CLEO III experiment, $\mathcal{B}[Y(3S) \rightarrow \gamma \eta_b] < 4.3 \times 10^{-4}$ at 90% confidence level, is based on 1.39 fb^{-1} of $Y(3S)$ data [4].

The data sample used in this study was collected with the *BABAR* detector [5] at the PEP-II asymmetric-energy e^+e^- storage rings. It consists of 28.0 fb^{-1} of integrated luminosity collected at a e^+e^- center-of-mass (c.m.) energy of 10.355 GeV, corresponding to the mass of the $Y(3S)$ resonance. Additional samples of 2.4 and 43.9 fb^{-1} were collected 30 MeV below the $Y(3S)$ [below- $Y(3S)$] and 40 MeV below the $Y(4S)$ [below- $Y(4S)$] resonances, respectively, and are used for background and calibration studies. The trajectories of charged particles are reconstructed using a combination of five layers of double-sided silicon strip detectors and a 40-layer drift chamber, all operated inside the 1.5-T magnetic field of a superconducting solenoid. Photons are detected using a CsI(Tl) electromagnetic calorimeter (EMC), which is also inside the coil. The energy resolution for photons varies from 2.9% (at 600 MeV) to 2.5% (at 1400 MeV).

The signal for $Y(3S) \rightarrow \gamma \eta_b$ is extracted from a fit to the inclusive photon energy spectrum in the c.m. frame. Any reference to photon energy hereafter will be in the c.m. frame, unless otherwise noted.

The monochromatic photon from the decay appears as a peak on top of a smooth nonpeaking background from continuum ($e^+e^- \rightarrow q\bar{q}$ with $q = u, d, s, c$) events and bottomonium decays. Two other processes produce peaks in the photon energy spectrum close to the signal region. Double radiative decays $Y(3S) \rightarrow \gamma \chi_{bJ}(2P)$, $\chi_{bJ}(2P) \rightarrow \gamma Y(1S)$, $J = 0, 1, 2$ produce a broad peak centered at 760 MeV due to photons from decays of the $\chi_{bJ}(2P)$ states. The peaks from the three $\chi_{bJ}(2P)$ transitions appear merged due to photon energy resolution and the Doppler broadening that arises from the motion of the $\chi_{bJ}(2P)$ in the c.m. frame. This $\chi_{bJ}(2P)$ photon peak is well separated from the signal region of interest (around $E_\gamma = 900$ MeV).

We use the peak as a tool to verify the optimization of the selection criteria and to determine signal reconstruction efficiencies and the absolute photon energy scale. The other process leading to a peak near 860 MeV in the photon energy spectrum is the radiative production of the $Y(1S)$ via initial state radiation (ISR) $e^+e^- \rightarrow \gamma_{\text{ISR}} Y(1S)$. Knowledge of the magnitude and photon energy line shape of this background is crucial in extracting the η_b signal.

We employ a simple set of selection criteria to suppress the backgrounds while retaining a high signal efficiency. Decays of the η_b via two gluons, expected to be a large component of its decay modes, have high track multiplicity. Hadronic events are selected by requiring four or more charged tracks in the event and that the ratio of the second to zeroth Fox-Wolfman moments [6] be less than 0.98.

Photon candidates are required to be isolated from all charged tracks. To ensure that their shapes are consistent with an electromagnetic shower, the lateral moments [7] are required to be less than 0.55. The signal photon candidate is required to lie in the central angular region of the EMC, $-0.762 < \cos(\theta_{\gamma, \text{LAB}}) < 0.890$, where $\theta_{\gamma, \text{LAB}}$ is the angle between the photon and the beam axis in the laboratory frame. This requirement ensures high reconstruction efficiency and good energy resolution, and reduces the contributions of ISR photons from $e^+e^- \rightarrow \gamma_{\text{ISR}} Y(1S)$ events.

Because of the fact that there is no preferred direction in the decay of the spin-zero η_b , the correlation of the direction of the photon momentum in the c.m. frame with the thrust axis [8] of the η_b is small. In contrast, there is a strong correlation between the photon direction and thrust axis in continuum events. The thrust axis is computed with all charged tracks and neutral calorimeter clusters in the event, with the exception of the signal photon candidate. We require $|\cos\theta_T| < 0.7$ to reduce continuum background, where θ_T is the angle between the thrust axis and the signal photon candidate in the c.m. frame.

Photons from π^0 decays are one of the main sources of background. A signal photon candidate is rejected if it combines with another photon in the event to form a π^0 candidate within 15 MeV/ c^2 of the nominal π^0 mass. To maintain high signal efficiency, we require the second photon of the π^0 candidate to have an energy in the laboratory frame greater than 50 MeV.

The above-mentioned selection criteria were chosen by optimizing the $S/\sqrt{B}$ ratio between the expected signal yield (S) and the background (B). The signal sample in the optimization is provided by a detailed Monte Carlo (MC) simulation [9]. Since no reliable event generators exist to simulate the background photon distribution, especially from bottomonium decays, a small fraction (9%) of the $Y(3S)$ data is used in the optimization to model the background in the region $0.85 < E_\gamma < 0.95$ GeV. To avoid potential bias, these data are not used in the final fit of the photon energy spectrum. This optimization procedure,

when applied to the $\chi_{bJ}(2P)$ yield in data in place of the simulated signal, yields the same optimal selection criteria. The final reconstruction efficiency evaluated from the simulated signal MC events is 37%.

The remaining $Y(3S)$ data used for the analysis has an integrated luminosity of 25.6 fb^{-1} , which corresponds to (109 ± 1) million $Y(3S)$ events.

To extract the η_b signal, we perform a binned maximum likelihood fit of the E_γ spectrum with $0.5 < E_\gamma < 1.1 \text{ GeV}$ with four components: nonpeaking background, $\chi_{bJ}(2P) \rightarrow \gamma Y(1S)$, $\gamma_{\text{ISR}} Y(1S)$, and the η_b signal.

The nonpeaking background is parametrized by the following probability density function (PDF), $f(E_\gamma) = A(C + \exp[-\alpha E_\gamma - \beta E_\gamma^2])$.

The form of the $\chi_{bJ}(2P)$ PDF is complicated by the presence of Doppler broadening. Crystal Ball (CB) functions [10] are used as phenomenological PDFs for the three $\chi_{bJ}(2P) \rightarrow \gamma Y(1S)$ shapes. The CB function is a Gaussian modified to have an extended, power-law tail on the low (left) side. The relative rates and peak positions of the $\chi_{bJ}(2P)$ components are fixed to their world-averaged (PDG) values [11]. The parameters describing the low-side tail of the CB function are common to all three of the $\chi_{bJ}(2P)$ peaks. The $\chi_{bJ}(2P)$ PDF parameters are determined by fitting the photon energy spectrum, with the signal region (840 to 960 MeV) excluded, after subtraction of the nonpeaking background. All of the $\chi_{bJ}(2P)$ PDF parameters from this fit, with the exception of the overall normalization, are fixed in the ultimate fit to the full photon energy spectrum.

The PDF of the peaking background from ISR $Y(1S)$ production is parametrized as a CB function form whose parameters are determined from simulated events. To estimate the rate of this continuum component in $Y(3S)$ data, we use the below- $Y(3S)$ and below- $Y(4S)$ data. Figure 1 shows the E_γ distribution in the below- $Y(4S)$ data, after applying the selection criteria and subtracting the non-

peaking background. The fit with a CB function yields $35\,800 \pm 1600$ events. Extrapolating the cross section to the $Y(3S)$ energy and correcting for the luminosity ratio and the small difference in detection efficiency, the ISR photon background contribution in the final analysis is estimated to be $25\,200 \pm 1700$ events. The error includes systematic uncertainties. This is consistent with and more precise than the rate estimated using the below- $Y(3S)$ data.

In the final fit of the whole E_γ distribution to extract the η_b signal, all parameters of the $\chi_{bJ}(2P)$ peak and the ISR $Y(1S)$ PDFs are fixed to the values from the fits described above.

The η_b signal PDF is a nonrelativistic Breit-Wigner function convolved with a CB function to account for the experimental E_γ resolution. The CB parameters are determined from signal MC simulations with the η_b width set to zero. Since the width of the η_b is not known, we have chosen a nominal value of $10 \text{ MeV}/c^2$ for the width. Theoretical predictions based on the expected ratio of the two-photon and two-gluon widths range from 4 to 20 MeV [12]. The free parameters in the fit are the η_b peak position and signal yield, the $\chi_{bJ}(2P)$ yield, and all of the nonpeaking background PDF parameters.

Figure 2(a) shows the photon energy spectrum and the fit result. The nonpeaking background is dominant with only the prominent $\chi_{bJ}(2P)$ peak visible. In Fig. 2(b) we show the detail of the signal region, after subtracting the nonpeaking background. The line shapes of the three peaking components, $\chi_{bJ}(2P)$, ISR $Y(1S)$, and the η_b signal are clearly visible. The χ^2 per degree of freedom from the fit is $147/113 = 1.3$. Finally Fig. 2(c) shows the data points with all components except the η_b signal subtracted, overlaid with the η_b signal PDF. The fitted η_b signal yield is $19\,200 \pm 2000 \pm 2100$ events, where the first error is statistical and the second systematic. A total systematic uncertainty of 11% is estimated by varying the Breit-Wigner width in the η_b PDF to 5, 15, and 20 MeV, setting the ISR $Y(1S)$ component to $\pm 1\sigma$ of the nominal rate, and varying the PDF parameters fixed in the fit by $\pm 1\sigma$. The largest contribution (10%) is from the η_b width variation.

The η_b signal significance is estimated using the ratio $\log(L_{\text{max}}/L_0)$, where L_{max} and L_0 are the likelihood values obtained from the nominal fit and from a fit with the η_b PDF removed, respectively. Fits have been performed where the parameters entering the systematic uncertainties have been varied within their errors. Data have then been fitted with all parameters simultaneously moved by 1 standard deviation in the direction of lower significance. This conservative approach yields a signal significance greater than 10 standard deviations.

As a cross check, we also perform a fit where the yield of the ISR $Y(1S)$ component is left free, and we obtain $24\,800 \pm 2300$ events for this component. This is consistent with the estimate using the below- $Y(4S)$ data and provides an important validation of the $\chi_{bJ}(2P)$ line shape

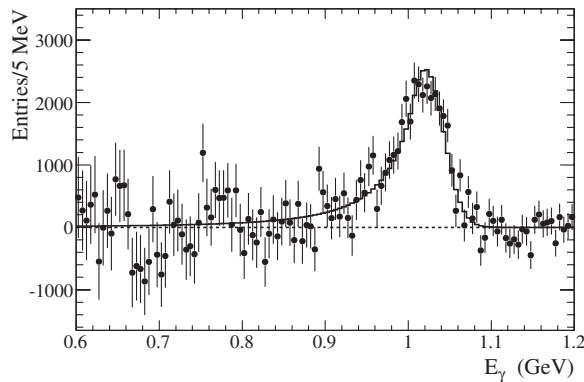


FIG. 1. Inclusive photon energy spectrum in the below- $Y(4S)$ data, with the nonpeaking background subtracted. The peak at 1.03 GeV is from the ISR process $e^+e^- \rightarrow \gamma_{\text{ISR}} Y(1S)$. The superimposed histogram corresponds to a fit with a CB function.

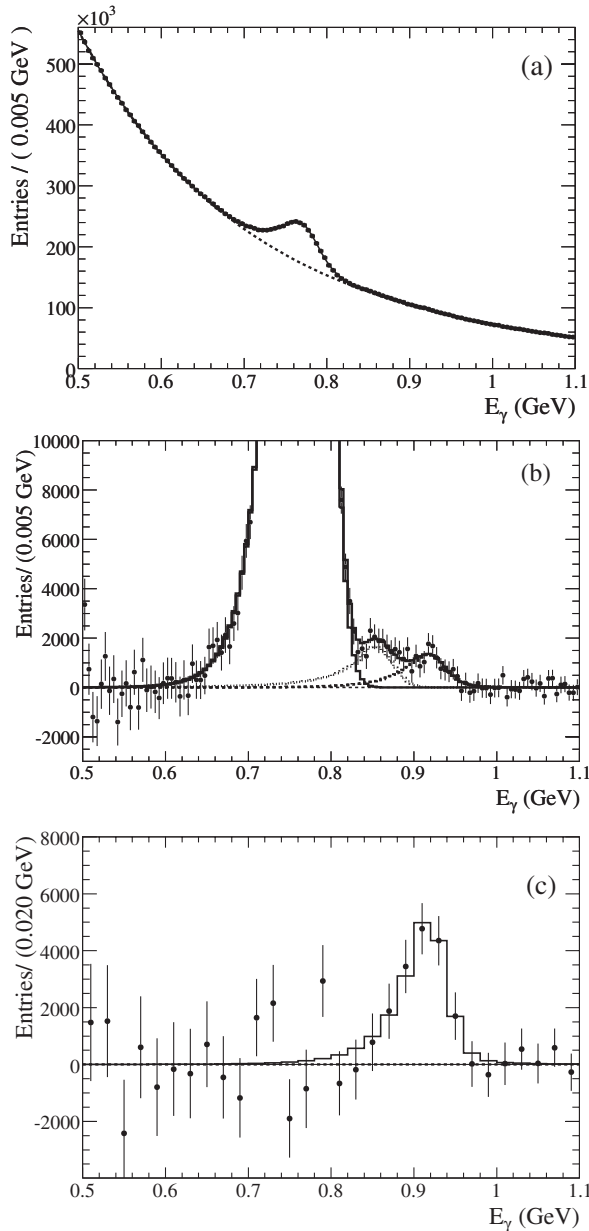


FIG. 2. (a) Inclusive photon spectrum in the region $0.50 < E_\gamma < 1.1$ GeV. The component PDFs determined from the fit are overlaid on the data points. A prominent $\chi_{bJ}(2P)$ peak is clearly seen. The dashed line corresponds to the nonpeaking background component. (b) Inclusive photon spectrum after subtracting the nonpeaking background, with PDFs for the $\chi_{bJ}(2P)$ peak (solid line), ISR $Y(1S)$ (dotted line), η_b signal (dashed line) and the sum of all three (solid line). (c) Inclusive photon spectrum after subtracting all components except the η_b signal. The CB function shape describes the data points well.

parametrization. The yield and peak position of the η_b signal from this fit are unchanged.

The E_γ signal peak value from the fit is $917.4^{+2.1}_{-2.8}$ MeV. We apply a photon energy calibration shift of 3.8 ± 2.0 MeV, obtained by comparing the fitted position of the $\chi_{bJ}(2P)$ peak to the known PDG value. After including

an additional systematic uncertainty of 1.3 MeV from the fit variations described above, we obtain a value of $E_\gamma = 921.2^{+2.1}_{-2.8} \pm 2.4$ MeV for the η_b signal.

The η_b mass derived from the E_γ signal is $M(\eta_b) = 9388.9^{+3.1}_{-2.3} \pm 2.7$ MeV/ c^2 . Using the PDG value of 9460.3 ± 0.3 MeV/ c^2 for the $Y(1S)$ mass, we determine the $Y(1S)$ - η_b mass splitting to be $71.4^{+2.3}_{-3.1} \pm 2.7$ MeV/ c^2 .

The value we measure for the splitting is larger than most predictions based on potential models [2], but reasonably in agreement with predictions from lattice calculations [13]. The mass splitting between the $Y(1S)$ and the $\eta_b(1S)$ is a key ingredient in many theoretical calculations. The precision of our measurement will allow, among others, a more precise determination of the lattice spacing [13] and new precision determinations of α_s [14].

We estimate the branching fraction by correcting the signal yield with the reconstruction efficiency (ϵ) from simulated signal MC events, and then dividing it by the number of $Y(3S)$ events in the data sample. The branching fraction of the decay $Y(3S) \rightarrow \gamma \eta_b$ is found to be $(4.8 \pm 0.5 \pm 1.2) \times 10^{-4}$, where the first uncertainty is statistical and the second systematic. The systematic uncertainty of 25% comes from uncertainties in the signal yield (11%) and ϵ (22%). The latter is obtained by comparing the yield of $\chi_{bJ}(2P)$ in data to the number of expected events, which is calculated from the known branching fractions [11], the number of $Y(3S)$ events, and MC reconstruction efficiency of $\chi_{bJ}(2P)$. They show a 13% discrepancy, but are consistent within the errors. We assign the full difference to the systematic uncertainty. A total uncertainty in ϵ is obtained, after adding the uncertainties in the $\chi_{bJ}(2P)$ branching fractions (18%).

In conclusion, we have observed, with a significance of 10 standard deviations, the radiative decay of the $Y(3S)$ to a narrow state lying slightly below the $Y(1S)$. The most likely interpretation of the signal peak is the $Y(3S)$ transition to the bottomonium ground state, although other hypotheses, such as a radiative transition to a light Higgs boson, are not excluded. Under the bottomonium interpretation, this is the first evidence for the η_b bottomonium state, the pseudoscalar partner of the $Y(1S)$. The mass of the η_b is $9388.9^{+3.1}_{-2.3} \pm 2.7$ MeV/ c^2 , which corresponds to a mass splitting between the $Y(1S)$ and the η_b of $71.4^{+2.3}_{-3.1} \pm 2.7$ MeV/ c^2 . The estimated branching fraction of the decay $Y(3S) \rightarrow \gamma \eta_b$ is found to be $(4.8 \pm 0.5 \pm 1.2) \times 10^{-4}$.

We are grateful for the excellent luminosity and machine conditions provided by our PEP-II colleagues, and for the substantial dedicated effort from the computing organizations that support BABAR. We thank Bob McElrath and Michael Peskin for helpful discussions. The collaborating institutions wish to thank SLAC for its support and kind hospitality. This work is supported by DOE and NSF (USA), NSERC (Canada), CEA and CNRS-IN2P3

(France), BMBF and DFG (Germany), INFN (Italy), FOM (The Netherlands), NFR (Norway), MES (Russia), MEC (Spain), and STFC (United Kingdom). Individuals have received support from the Marie Curie EIF (European Union) and the A. P. Sloan Foundation.

*Deceased.

[†]Present and permanent address: Temple University, Philadelphia, PA 19122, USA.

[‡]Present and permanent address: Tel Aviv University, Tel Aviv, 69978, Israel.

[§]Also at Università di Perugia, Dipartimento di Fisica, Perugia, Italy.

^{||}Also at Università di Roma La Sapienza, I-00185 Roma, Italy.

[¶]Present and permanent address: University of South Alabama, Mobile, AL 36688, USA.

**Also at Università di Sassari, Sassari, Italy.

- [1] S. W. Herb *et al.*, Phys. Rev. Lett. **39**, 252 (1977); W. R. Innes *et al.*, Phys. Rev. Lett. **39**, 1240 (1977); **39**, 1640(E) (1977).
- [2] For a comprehensive review, see N. Brambilla *et al.* (Quarkonium Working Group), CERN Yellow Report, No. CERN-2005-005 (2005).
- [3] S. Godfrey and J. L. Rosner, Phys. Rev. D **64**, 074011 (2001); **65**, 039901(E) (2002), and references therein.
- [4] M. Artuso *et al.* (CLEO III Collaboration), Phys. Rev. Lett. **94**, 032001 (2005).
- [5] B. Aubert *et al.* (BABAR Collaboration), Nucl. Instrum. Methods Phys. Res., Sect. A **479**, 1 (2002).
- [6] G. C. Fox and S. Wolfram, Nucl. Phys. **B149**, 413 (1979).
- [7] A. Drescher *et al.* (ARGUS Collaboration), Nucl. Instrum. Methods Phys. Res., Sect. A **237**, 464 (1985).
- [8] S. Brandt *et al.*, Phys. Lett. **12**, 57 (1964); E. Fahri, Phys. Rev. Lett. **39**, 1587 (1977).
- [9] S. Agostinelli *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **506**, 250 (2003); T. Sjöstrand and M. Bengtsson, Comput. Phys. Commun. **43**, 367 (1987).
- [10] J. E. Gaiser, Ph.D. thesis, Stanford University [SLAC-R-255] (1982).
- [11] W.-M. Yao *et al.* (Particle Data Group), J. Phys. G **33**, 1 (2006) and 2007 partial update for the 2008 edition.
- [12] W. Kwong *et al.*, Phys. Rev. D **37**, 3210 (1988); C. S. Kim, T. Lee, and G. L. Wang, Phys. Lett. B **606**, 323 (2005); J. P. Lansberg and T. N. Pham, Phys. Rev. D **75**, 017501 (2007).
- [13] A. Gray *et al.* (HPQCD and UKQCD Collaborations), Phys. Rev. D **72**, 094507 (2005); T. Burch and C. Ehmann, Nucl. Phys. **A797**, 33 (2007); T. W. Chiu *et al.* (TWQCD Collaboration) Phys. Lett. B **651**, 171 (2007).
- [14] B. A. Kniehl, A. A. Penin, A. Pineda, V. A. Smirnov, and M. Steinhauser, Phys. Rev. Lett. **92**, 242001 (2004).